

# **The Automated Satellite Data Processing System**

## **AVHRR Processing**

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# **The Automated Satellite Data Processing System: AVHRR Processing**

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# **Part I. Advanced Very High Resolution Radiometer**

The chapters in Part I form a User's Guide for the processing of the Advanced Very High Resolution Radiometer (AVHRR) within the Automated Processing System.

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# Introduction

The Advanced Very High Resolution Radiometer (AVHRR) is a space-borne sensor embarked on the National Oceanic and Atmospheric Administration (NOAA) family of polar orbiting platforms (POES). AVHRR instruments measure the reflectance of the Earth in up to 6 relatively wide (by today's standards) spectral bands. The first two bands are centred around the red (0.6 micrometer) and near-infrared (0.9 micrometer) regions. In the short-wave-infrared regions two bands sample the 1.6 and 3.8 micrometer regions. These bands are given the designations 3A and 3B. The last two sample the thermal radiation emitted by the planet, around 11 and 12 micrometers, respectively. They are commonly designated band 4 and 5.

The first AVHRR instrument actually was a 4-channel radiometer (bands 1, 2, 3B, and 4). The second version of the AVHRR instrument added the 12 micrometer band. The latest version (known as AVHRR/3, first carried on the NOAA-15 platform launched in May 1998) acquires data in a 6th channel located at 1.6 micrometer. Due to the original design of the transmission format which allowed for only five bands, the new SWIR band replaces the original 3.8 micrometer band during day time viewing.

The primary purpose of these instruments is to monitor clouds and to measure the thermal emission (cooling) of the Earth. These sensors have proven useful for a number of other applications, however, including the surveillance of land surfaces, ocean state, aerosols, etc. AVHRR data are particularly relevant to study climate change and environmental degradation because of the comparatively long records of data already accumulated (over 20 years). The main difficulty associated with these investigations is to properly deal with the many limitations of these instruments, especially in the early period (sensor calibration, orbital drift, limited spectral and directional sampling, etc).

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# Part II. Command Line Reference

The chapters in Part II form a reference guide for each program available in the Automated Processing System for processing AVHRR data.

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# Name

avhArea -- determines the file extents of a NESDIS Level-1B data file which covers an image map.

avhArea

avhArea [*options*] *mapname filename*

## Description

Determines the file extents (start/stop pixel/line) of a NESDIS Level-1B file (still in sensor projection) that covers a map.

The command **avhArea** begins by reading in the map from the mapfile. If the file can not be opened or the named map is not in the file, a diagnostic is printed and the program will exit.

Next, the AVHRR file is opened and the navigation information initialized. If unable to open the file or get the navigation information from the file, the program will print a diagnostic and exit.

Once the navigation has been set, **avhArea** reads in every scan line and determines the latitude and longitude for 82 pixels over the entire scan line (or approximately every 25th pixel). For each point that falls within the desired maps, the starting and stopping sample (or column) number of the file is determined. The line extents are also determined by the first line that contains data that falls within the box and the last line that falls outside the box again. The file extents are adjusted to be slightly larger than those found by the above procedure to ensure that no data within the region is missed. These file extents will be printed to the screen. These are printed to stdout: starting pixel, space, ending pixel, space, starting line, space, ending line.

If the entire file covers the image map, then "Complete coverage" will be written to stdout. If no part of the file covers the image map, then "No coverage" will be written to stdout.

Based on the landmask, **avhArea** can also determine if any pixels within the region fell over water. If not samples fell over water then the message "No Water Coverage" is added. This can be used to determine if the file is to be processed even when it covers the interested area.

## Options

-f type

Set the format type of the input AVHRR file. Valid responds are 1 for NESDIS Level-1B format (NOAA-6 to NOAA-14), 2 for NESDIS KLM format (NOAA-15 to NOAA-16), 3 for NESDIS Level-0 format, 4 for NRL Terascan HDF format

-l

Don't output start/stop line locations

-L file

Use file as the input land mask file. Defaults to \$APS\_DATA/landmask.dat

-M mapFile

Use the given map file to find mapName. Defaults to \$APS\_DATA/maps.hdf

- n #  
Set the number of points across and down the image used to search for data coverage. The default is 82 points which yields a control point roughly every 25 pixels and lines. For small regions - that might "fall between the cracks" - this can be set to a higher number to create a finer grid.
- p  
Don't output start/stop pixel locations
- v  
Verbose output
- help  
Display program help.
- version  
Display program name version and time of compilation.

## Environmental Variables

APS\_DATA  
The location of the APS data directory.

## Examples

The examples below show the same input file run against two different geographical areas. The last examples shows the result of trying to use an invalid input.

### Example 1. Use of avhArea

```
$ avhArea GulfOfMexico noaa-14.970205200124.lac
1 1638 1021 2551
$ avhArea -p -M my_maps.hdf GulfOfMexico noaa-14.970205200124.lac
1021 2551
$ avhArea EastSea noaa-14.970205200124.lac
No coverage
$ avhArea Junk noaa-14.970205200124.lac
Map (Junk) does not exist in file ($APS_DATA/maps.hdf).
$ echo $?
1
```

---

# Name

avhClouds -- produces a bit image of tests used to determine cloud contaminated pixels

```
avhClouds
avhClouds in.hdf [out.hdf]
```

## Description

**avhClouds** will determine the cloud cover using several different tests based on the cloud detection algorithm of Saunder and Kriebel (1988). These tests require brightness temperatures from channels 3, 4, and 5. If the scene is a daytime scene, then the albedo from channels 1 and 2 are also required. Each test will set a bit in the output file. The tests are divided into day and night tests as well as land/sea/coast test.

The input file should be the output from avhIngest(1). That program produces the calibrated radiances and brightness temperature products as well as the `l2_flags` product used by this **avhClouds**.

These tests are performed on a 3x3 box with all edge pixels being marked as cloud automatically. The input file is read in and used to determine on a pixel by pixel basis some of the characteristics needed by the algorithm for that pixel.

If all the surrounding pixels are determined to be land, the center pixel is marked as land. If all the surrounding pixels are determined to be sea (not land), the center pixel is marked as sea. Otherwise, the center pixel is marked as coast. The program **avhIngest** uses a landmask file to set the LAND flag (bit 2).

For the day or night characteristic of a pixel, **avhClouds** will set all pixels to day or night, if upon initialization, the code determines that all pixels are day or night. This is determined by examining the Sun elevation for the four corner points and center of the input image. If all five points are defined as day, then all pixels are marked as day. If all four points are defined as night, then all pixels are marked as night. A Sun elevation greater than 15 degrees implies day and less than or equal to 15 implies night.

If this gross day/night check fails, then the DAY\_TIME flag set in the `l2_flags` product in the input file by **avhIngest** will be consulted. In a similar manner to land/sea/coast, the DAY\_TIME flag (bit 3) for surrounding pixels is examined to determine if the center pixel is DAY or NIGHT. If all surrounding pixels are defined as DAY\_TIME, then the center pixel is marked as DAY. If all surrounding pixels are defined as NIGHT\_TIME, then the center pixel is marked as NIGHT.

## Channel 4 Gross Cloud Check

If channel 4 temperature is too low, it is assumed that these are cloud-top temperatures. All values over sea that are less than 273.15 degrees Kelvin (0 degrees Celsius) are flagged as clouds. For land and coast pixels, 263.15 degrees Kelvin (-10 degrees Celsius) are used. Bit 16 (CLD\_CH4\_GROSS\_CLOUD\_CHECK) represents this test.

## Channel 4 Spatial Coherence Test

The standard deviation in a 3x3 box surrounding the pixel in question for channel 4 is determined. If the standard deviation for sea (day or night) or land (night only) are greater than the thresholds (0.35 and 1.75), then the pixel is masked as cloudy. Bit 17 (CLD\_CH4\_SPATIAL\_COHERENCE) represents this test.

## Visible Channel Test

During the day, a high value of in channel 2 may indicate clouds cover. For land pixels, channel 1 is used for the test if available. The threshold of these land pixels is 40.0. For sea pixels, the visible threshold is 10. For coastal pixels, the visible threshold is 15.0. Bit 18 (CLD\_CH2\_GROSS\_CHECK) represents this test.

## Channel 2 Spatial Coherence Test

For a day pixel, the standard deviation in a 3x3 box surrounding the pixel in question of channel 2 is determined. If the deviation for sea (day or night) are greater than 0.4, then the pixel is masked as cloudy. Bit 19 (CLD\_CH2\_SPATIAL\_COHERENCE) represents this test.

## NIR/VIS Test

During the day, the ratio of the NIR (channel 2) over VIS (channel 1) may indicate clouds. For land pixels, the default threshold is 0.0. For sea pixels the default threshold is 0.75. Bit 20 (CLD\_NIR\_VIS) represents this test.

## Low Fog and Uniform Stratus Check

For a night pixel, the channel 4 - channel 3 difference greater than a given threshold (default 1.0) indicates low fog or uniform stratus clouds. For these pixels, bit 21 (CLD\_LOW\_FOG\_UNIFORM\_STRATUS) represents this test.

## Medium and High Cloud Check

If the pixel in question is a night pixel, then the difference in ch3 - ch5 will indicate medium and/or high clouds. The default threshold for this value is 1.5. Bit 22 (CLD\_MEDIUM\_HIGH\_CLOUD) represents this test.

## Thin cirrus cloud check

If the pixel in question is a night pixel, then the difference in ch4 - ch5 is consulted against a 2-D table that will vary based on ch4 and the satellite zenith angle. If the channel 4-5 difference is greater than the found threshold, the pixel is flagged as clouds. Bit 23 (CLD\_THIN\_CIRRUS) represents this test.

## Options

- T file  
This runs the program in 'trace' mode.
- v  
Verbose mode.
- help  
Display program help.
- version  
Display program name version and time of compilation.

## Reference

Saunders, R.W. and Kriebel, K.T., 1988, *An improved method for detecting clear sky and cloudy radiances from AVHRR data* Int. Journal of Remote Sensing, 1988, Vol 9, No. 1, pp 123-150.

---

# Name

avhDump -- dumps AVHRR video data from a NESDIS Level-1B file.

```
avhDump  
avhDump -c n [-f 1] filename chan.bin
```

# Description

**avhDump** is used to dump AVHRR video data from a NESDIS Level-1B file. Currently, it only supports the dumping of AVHRR video data which must be selected with the `-c` option.

The output file is written as 16-bit integers in a flat binary format which has 2048 (LAC/HRPT) or 409 (GAC) columns by *n* number of rows. The number of columns and rows are printed by this program. Also, if the user knows the file type, the number of rows is therefore known and the number of columns can be computed by dividing the size of the file by 2 times the number of columns.

# Options

- c n  
Select a channel to output. Must be between 1 and 5.
- f n  
Select format of the input file. Use 1 for NESDIS Level-1B, 2 for NESDIS KLM Level-1B, 3 for NESDIS HRPT Level-0, and 4 for NRL Terascan HDF format.
- v  
Turn on verbose mode.
- help  
Display program help.
- version  
Display program name version and time of compilation.

---

# Name

avhImage -- creates a simple graphics image file from a NESDIS Level-1B file.

avhImage

avhImage *[[options]] filename image.ext*

## Description

**avhImage** is used to make a quick image from an AVHRR data file. By default, the program will read channel four counts (10-bits) and shift two bits to the right two for an output of the top 8 bits. If the input file is a LAC or HRPT file every fourth line and sample are written. If the input file is a GAC file, then every line/sample are written to the file.

The output file may be written as a PNM grayscale raw file, TIFF, PNG, or SGI RGB file depending on the compilation of the program. The available types may be obtained by running **avhImage --help**. The type is selected based on the extension of the output file.

## Options

-c n

Select a channel to output. Must be between 1 and 5.

-F

Full output. If the input file is a LAC/HRPT file, this option writes the entire image to a file.

-t type

Set the format type of the output image file. Valid responds are based on the compilation of the program. The *type* given is given as an image "extension". For example, *tiff*.

-f type

Set the format type of the input AVHRR file. Valid responds are 1 for NESDIS Level-1B format (NOAA-6 to NOAA-14), 2 for NESDIS KLM format (NOAA-15 to NOAA-16), 3 for NESDIS Level-0 format, 4 for NRL Terascan HDF format

--help

Display program help.

--version

Display program name version and time of compilation.

## Examples

To determine which formats are available, we run **avhImage** using the **--help** option.

```
$ avhImage --help
```

```
Usage: avhImage [OPTION] INPUT OUTPUT
```

Creates a subsampled grayscale image of channel 4 from the AVHRR data stored in a NESDIS Level-1B file, KLM Level-1B file, or an HRPT Level-0 file.

### OPTIONS

```
-c n      use channel n for output
-f type   set input file type
          1=NESDIS Level-1B file
          2=NESDIS KLM Level-1B file
          3=NESDIS Level-0 file
          4=TeraScan HDF file
-F        do full output
-t type   set output file type, see below
--help   this output
--version version information
```

INPUT must be a one of the formats above  
OUTPUT is a graphics file specified by its extension. It can one of the following:

```
pnm      for PNM Format
png      for PNG Format
tiff     for TIFF Format
```

Since this version allows us to create PNG files, we create a quick-look PNG image of our input file using the following command:

```
$ avhImage NSS.HRPT.NL.D02142.S2044.E2056.B0858282.MO chan4.png
```

---

# Name

avhInfo -- queries information about a NESDIS Level-1B file(s).

```
avhInfo
avhInfo file1 file2 [file3]
avhInfo option file
```

# Description

Run without options, **avhInfo** will write a report for each input file indicating satellite id, data type, etc. It may also be run with a single option and print the input file(s) value for that option. The first method is intended for interactive use at the shell prompt and the second method is intended for use within a shell program.

# Options

-year	4-digit year of input file.
-doy	3-digit day of year of input file.
-month	3-character month of input file. Months are 'jan', 'feb', 'mar', 'apr', 'may', 'jun', 'jul', 'aug', 'sep', 'oct', 'nov', 'dec'
-time	6-digit time (HHMMSS) of input file.
-hour	2-digit hour (HHMMSS) of input file.
-min	2-digit min (MM) of input file.
-sec	2-digit second (SS) of input file.
-start_time	start time of input file.
-end_time	end time of input file.
-dsn	The NOAA defined Data Set Name from the TBM/ARS header.
-type	1-digit code for datatype, where: 1=LAC, 2=GAC, 3=HRPT
-sat	3-character satellite name. Names are 't-n', 'n06', 'n07', 'n08', 'n09', 'n10', 'n11', 'n12', 'n14', 'n15', or 'n16'.

- `-sat_code` 3-character satellite name. Names are 'TN', 'NA', 'NC', 'NE', 'NF', 'NG', 'ND', 'NJ', 'NK', or 'NL'.
- `-name` Generate a file name in the following format as SSS.YYYY.MMDD.HHMM.T, where T is 'l' for LAC, 'h' is for HRPT and 'g' is for GAC.
- `--help` Display program help.
- `--version` Display program name version and time of compilation.

## Examples

Here is how a Bourne shell script function might use **avhInfo** to set the name of the output filenames:

### Example 2. Use of avhInfo within shell

```
set_name()  
{  
    sat='avhInfo -sat $1'  
    yr='avhInfo -year $1'  
    jday='avhInfo -doy $1'  
    time='avhInfo -time $1'  
    file=$sat.$yr$jday.$time.llb  
}
```

Here is an interactive use of **avhInfo**:

### Example 3. Interactive use of avhInfo

```
$ avhInfo *.hrpt
Filename:n950622.2233.n14.hrpt
Format:NESDIS
Starting Time:06/22/95 22:32, 173
Ending Time:06/22/95 22:44, 173
Satellite:n14
Datatype:LAC (Local Area Coverage)
Total Scans:4012 with no gaps
```

```
Filename:n950623.0256.n12.hrpt
Format:NESDIS
Starting Time:06/23/95 02:56, 174
Ending Time:06/23/95 03:08, 174
Satellite:n12
Datatype:LAC (Local Area Coverage)
Total Scans:4331 with no gaps
```

---

# Name

avhIngest -- produces calibrated albedo and brightness temperature image from AVHRR data.

avhIngest

avhIngest *avhrr.1b out.hdf [channels]*

# Description

This program reads the AVHRR data from a NESDIS Level-1B file, NESDIS KLM Level-1B, or NESDIS Level-0, or NRL TeraScan HDF Level-0 file and writes out calibrated data to an APS HDF file format. The albedo data is calibrated using the slope and intercept specified in the input file. The brightness temperatures are calibrated using an algorithm by Brown or by the calibration coefficients specified in the input file for KLM.

The user may select those parameters in the following table marked with a dagger on the command line. By default all of these parameters will be created unless unavailable due to the input.

## Table 1. AVHRR Ocean Parameters

<b>Product</b>	<b>Description</b>
l2_flags	Level-2 Flags image
albedo_ch1†	Percent albedo image for visible channel 1
albedo_ch1†	Percent albedo image for visible channel 2
btemp_ch3†	Brightness energy temperature (deg Kelvin) image for IR channel 3
btemp_ch4†	Brightness energy temperature (deg Kelvin) image for IR channel 4
btemp_ch5†	Brightness energy temperature (deg Kelvin) image for IR channel 5
albedo_ch3†	Percent albedo image for visible channel 3A
latitudes†	Latitudes
longitudes†	Longitudes
solz†	Solar zenith angles
sola†	Solar azmiuth angles
senz†	Sensor zenith angles
sena†	Sensor azmiuth angles
secsat†	Secant of the view angle

## Options

-a angle

If *angle* is defined then it is used to reduce the swath of the input image. This option can only be used for LAC/HRPT files and will calculate the number of pixels to reduce the image. It can be used to prevent the large pixels from the edge of the swath to be output. If angle is less than 1.1, then it is assumed to be given in radians. Otherwise it is give in degrees. A negative angle will be converted to a positive one.

-B *isp iep isl iel irp irl*

These set up the subsection of the NESDIS Level-1B file to extract the data from. *isp* is the starting pixel number (1 to 2048). *iep* is the ending pixel number (1 to 2048, greater than *isp*). *isl* is the starting line (1 to *n*). *iel* is the ending line (1 to *n*, greater than *isl*). *irp* is the pixel subsampling factor. *irl* is the line subsampling factor.

-d

debug output

-L filename

use filename for land mask

-N

force day/night flag to night

-v

verbose output

--help

display program help

--version

Display program name version and time of compilation.

## Examples

This example will produce floating point outputs of albedo channels 1 and 2 for input to the turbidity program **avhTurbid**.

### Example 4. Use of avhIngest

avhIngest

---

```
$ avhIngest -1 R -2 R NJ1999360192854.L1B_HNAV avhrr.spk
$ ls -l avhrr*
-rw-r----- 1 aps    aps      1837 Jan  6 10:45 avhrr.avh
-rw-r----- 1 aps    aps     252020 Jan  6 10:45 avhrr.ctl
-rw-r----- 1 aps    aps    16671232 Jan  6 10:45 avhrr1.spk
-rw-r----- 1 aps    aps    16671232 Jan  6 10:45 avhrr2.spk
$ more *.avh
avhIngest 2.4 12/07/1999.
```

-----

Input Parameter Information

-----

Information from Header

Satellite ID: 9  
Start Time (YY DDD HH:MM:SS) : 1999 360 19:29:18  
Start Time (YY DDD HH:MM:SS) : 1999 360 19:34:57  
Datatype: LAC  
Tip Source: Embedded TIP  
Scene start/end samples numbers : 1/2048 2048  
Scene start/end scan line numbers : 1/2035 2035  
Sample/scan line reduction factors: 1/1

Start loop.

Processing routine: brown version 1.0  
Brown, et al, JGR Vol 98, c10:18257-68  
noaa-14 coefficients  
New Channel 1 slope/int 0.108100 -3.864800  
New Channel 2 slope/int 0.109000 -3.674900

Information about Output Files Created

-----

Percent albedo image for channel 1: avhrr1.spk  
Percent albedo image for channel 2: avhrr2.spk  
Brightness energy temperature (deg C) image for channel 3: not requested.  
Brightness energy temperature (deg C) image for channel 4: not requested.  
Brightness energy temperature (deg C) image for channel 5: not requested.

Input slope(s)/intercept(s) used in: DATA = (GRAY/SLOPE) + INTCP  
channel 1 (A1): Data output as reals.  
channel 2 (A2): Data output as reals.

Minimum and Maximum values for each channel processed:  
channel 1: -3.864800 105.316197  
channel 2: -3.674900 99.657097

Number of data gap scan lines encountered: 0  
Output image sizes (pixels/lines): 2048/2035  
Start/end pixels of data in image: 0/2047  
Start/end lines of data in image: 0/2034

Navigation Information for Output Images

-----

Control point file name: avhrr.ctl  
Corner latitudes/longitudes: \_\_\_\_\_  
Top : 30.440039/-45.958789 25.621875/-76.079688  
Bottom: 25.822852/-75.333203 44.229687/-86.860938  
(Satellite descending.)  
Minimum/maximum latitudes: 25.621875/50.522852  
Minimum/maximum longitudes: -86.860938/-45.958789

## Caveats

Presently this software can only handle NOAA satellites 12, 14, 15, 16, and 17.

## References

Kidwell, K, 1999. *KLM User's Guide*. NCDC/NESDIS, National Climatic Data Center, Washington, D.C.

Kidwell, K, 1991. *NOAA Polar Orbiter User's Guide*. NCDC/NESDIS, National Climatic Data Center, Washington, D.C.

Brown, J. W., O. B. Brown, and R. H. Evans, 1993. *Calibration of AVHRR infrared channels: a new approach to non-linear correction*. J. Geophys. Res. 98 (NC10), 18257-18268.

---

# Name

avhSST -- produces sea surface temperature images from brightness temperature images

avhSST

avhSST *in.hdf out.hdf*

# Description

This program will read brightness temperature data using channels 3, 4, and 5 produced by the program avhIngest(1) and generate an image file of sea surface temperature estimates. The algorithm is automatically selected based on satellite and time. The algorithms are from NOAA.

<b>SATELLITE</b>	<b>DAY ALGORITHM</b>	<b>NIGHT ALGORITHM</b>
NOAA-16	MCSST Split (4/5)	MCSST Split (4/5)
NOAA-15	NLSST Split (4/5)	NLSST Split (4/5)
NOAA-14	NLSST Split (4/5)	NLSST Triple (3/4/5)
NOAA-12	NLSST Split (4/5)	NLSST Split (4/5)
NOAA-11	NLSST Split (4/5)	NLSST Triple (3/4/5)

The day or night algorithm is based on whether the elevation of the sun is greater than 20 degrees at the center of the image.

# Options

- C                                Sets the output temperature scale to Celsius
- d                                Force use of day algorithm
- e number                        Select a specific equation
- F                                Sets the output temperature scale to Fahrenheit
- n                                Force use of night algorithm
- T file                            This runs the program in 'trace' mode
- v                                Verbose output

--help

Display program help

--version

Display program name version and time of compilation.

## References

The coefficients for the various SST algorithms were obtained by NOAAASIS. <http://noaasis.noaa.gov/NOAASIS/ml/sst.html>

---

# Name

avhScan -- dumps scan line information from a NESDIS Level-1B file

avhScan

avhScan [-f *type*] [-n *num*] *file target*

# Description

This program reads a NESDIS Level-1B file and writes to stdout information for each scan line based on the user specified target. The target can be one of the following: *vis-cal*, *latlon*, *prt*, *solar*, *telm*, *time*.

If the user selects *vis-cal*, then the calibration values (slope/intercept) for all five channels will be dumped.

If target *latlon* is selected, then **avhScan** will dump the scan line number, ascending/descending flag, and the number of valid lat/lon pairs followed by the 1st, 26th, and 51st lat/lon pair.

If *prt* is selected, then **avhScan** will dump the three prt counts stored in each scan line. Generally, these are duplicates of each other and contain the multiplexed PRT counts for each of the four PRTs. This dump, however, does not de-multiplex them.

If target *solar* is selected, then **avhScan** will dump the scan line number, and the number of meaningful solar zenith angles followed by the 1st, 26th, and 51st solar zenith angle.

Target *telm* will cause **avhScan** to dump the the five ramp calibration counts, the three PRT counts, the ten internal target view counts for all three channels, and the ten space view counts for all five channels.

If the user selects *time*, then **avhScan** will dump the time embedded in each scan line.

# Options

-f *type*

set input file type

1=NESDIS Level-1B file 2=NESDIS KLM Level-1B file 3=NESDIS Level-0 file

-n *n*

skip every nth record

## Examples

Each of the examples below are from the file noaa-14.970205200124.lac. They do not represent the full output of the program, but an excerpt to show the format used.

### Example 5. Use of avhScan

avhScan

```
$ avhScan -n20 noaa-14.970205200124.lac cal | more
 1  1  0.1081 -3.8648  0.1090 -3.6749  -0.0017  1.6917  -0.1673↵
159.7771 -0.1834 178.0051
 21 21  0.1081 -3.8648  0.1090 -3.6749  -0.0017  1.6917  -0.1673↵
159.7771 -0.1834 178.0051
 41 41  0.1081 -3.8648  0.1090 -3.6749  -0.0017  1.6917  -0.1673↵
159.7771 -0.1834 178.0051
 61 61  0.1081 -3.8648  0.1090 -3.6749  -0.0017  1.6603  -0.1664↵
159.3572 -0.1824 177.6636
 81 81  0.1081 -3.8648  0.1090 -3.6749  -0.0017  1.6603  -0.1664↵
159.3572 -0.1824 177.6636
101 101  0.1081 -3.8648  0.1090 -3.6749  -0.0017  1.6603  -0.1664↵
159.3572 -0.1824 177.6636
121 121  0.1081 -3.8648  0.1090 -3.6749  -0.0017  1.6408  -0.1654↵
159.1567 -0.1827 177.7323
141 141  0.1081 -3.8648  0.1090 -3.6749  -0.0017  1.6408  -0.1654↵
159.1567 -0.1827 177.7323
161 161  0.1081 -3.8648  0.1090 -3.6749  -0.0016  1.6267  -0.1638↵
158.3915 -0.1818 177.5354
181 181  0.1081 -3.8648  0.1090 -3.6749  -0.0016  1.6267  -0.1638↵
158.3915 -0.1818 177.5354
201 201  0.1081 -3.8648  0.1090 -3.6749  -0.0016  1.6267  -0.1638↵
158.3915 -0.1818 177.5354
221 221  0.1081 -3.8648  0.1090 -3.6749  -0.0016  1.6266  -0.1639↵
158.4392 -0.1818 177.5393
```

```
$ avhScan -n 50 noaa-14.970205200124.lac latlon
 1  1 051 0  8.56/ -77.41  6.76/ -89.91  4.64/-102.32
 51 51 000 0  8.56/ -77.41  6.76/ -89.91  4.64/-102.32
101 101 051 0  9.79/ -77.66  8.01/ -90.20  5.85/-102.63
151 151 051 0 10.30/ -77.76  8.54/ -90.32  6.37/-102.77
201 201 051 0 10.79/ -77.85  9.03/ -90.43  6.85/-102.89
```

```
$ avhScan noaa-14.970205200124.lac prt | more
 1  1  917  939  3
 2  2  917  683  3
 3  3  917  939  3
 4  4  917  811  3
 5  5  917  939  3
 6  6  917  811  3
```

```
$ avhScan -n20 noaa-14.970205200124.lac solar | more
 1  1 51  46.50 35.50 25.50
 21 21 51  47.00 36.00 26.00
 41 41 0  47.00 36.00 26.00
 61 61 51  47.00 36.00 26.50
 81 81 51  47.00 36.00 26.50
101 101 51  47.00 36.50 26.50
121 121 51  47.50 36.50 26.50
141 141 51  47.50 36.50 27.00
161 161 51  47.50 36.50 27.00
181 181 51  47.50 36.50 27.00
201 201 51  47.50 37.00 27.50
221 221 51  47.50 37.00 27.50
241 241 51  48.00 37.00 27.50
```

```
$ avhScan -n20 noaa-14.970205200124.lac telm | more
Record/Line 1/ 1
Ramp 644 367 860 413 527
```

---

## Name

avhScripts -- standard AVHRR processing scripts

avhScripts

## Description

avhScripts provides Bourne shell scripting functions to process AVHRR data from a NESDIS Level-1B file. The avhProcess is the main user interface and provides for the standard processing steps. These steps are:

- 
- 1        **Verify input file  
is NESDIS  
Level-1B file  
using filefmt  
program.**
  - 2        Determine if file  
covers user  
defined map  
using avhArea  
program.
  - 3        Get time  
information from  
file using avhInfo  
program.
  - 4        Parse  
*\$L3ProdList* to  
determine which  
intermediate  
products are  
required.
  - 5        Ingest AVHRR  
data and calibrate  
it using avhIngest  
program.
  - 6        If user selected  
"sst" as a product,  
generate it from  
brightness  
temperatures  
using avhSST  
program.
  - 7        Generate the  
cloud masks and  
add to "l2\_flags"  
using avhClouds  
program.
  - 8        If user selected  
any of the visible  
data products and  
input is day  
scene, generate  
them using the  
avhTurbid  
program.
  - 9        Warp all the  
products  
generated above  
to the defined  
map projection.
  - 10       For a large  
product array, tile  
and compress the  
data.
  - 11       If user has  
defined  

---

*\$AvhPreBrowse*<sub>25</sub>  
it is now called.  
(Normally, not  
defined.)
  - 12       If user had

## Usage

A minimalist executable script for processing AVHRR data must source both the `apsScripts` and `avhScripts` located in the APS bin directory, define two required variables (see `REQUIRED VARIABLES`), and call the script function `avhProcess` passing it the name of the NESDIS Level-1B file. Many other variables can be optionally set to modify the "normal" mode of operations. These must be set prior to the call to `avhProcess`.

```
#!/bin/bash
. $APS_DIR/lib/aps/apsScripts.sh
. $APS_DIR/lib/aps/avhScripts.sh
MapName=GulfOfMexico
MapExt=GOM
avhProcess $1 $0
```

This script must be placed in the `$AREAS_PROC` directory which is normally the directory *areas* located in the main APS directory. The script must have execution permissions.

## Required Variables

These variables are required to process an area. They provide the script an area to process. Most of the remaining variables have defaults which can be overridden. They are described in the next section.

### *MapName*

This is the name of the image map stored in the file `$MapFile`.

## Optional Variables

The following sections are variables that have defaults which the user can override to change the behaviour of the default processing. They are grouped together by subject.

### Product Selection

*L3ProdList*

This is a space delimited list of products to be written to the output data base. The following products are available:

<b>albedo_ch1</b>	<b>Percent albedo from channel 1</b>
albedo_ch2	Percent albedo from channel 2
albedo_ch3	Percent albedo from channel 3A
btemp_ch3	Brightness temperature from channel 3
btemp_ch4	Brightness temperature from channel 4
btemp_ch5	Brightness temperature from channel 5
sst	Sea surface temperature
c_660	Beam attenuation at 660 nm.
K_PAR	Diffuse Attenuation for photosynthetically active radiation
suspend	Total suspended solids concentration
Ray_ch1	Rayleigh reflectance for channel 1
ref_dif	Reflectance difference between channel 1 and channel 2

If not defined, the default value is "sst clouds c\_660". For example, the user might put the line

```
L3ProdList="btemp_ch3 btemp_ch4 btemp_ch5 sst"
```

to retain only the thermal channel information.

*\${prod}\_Scaling*

This option is currently only defined for albedo\_ch[1,2,3], btemp\_ch[3-5] and sst products. It defines the user's desired output scaling. For example, to produce an SST byte image with a scale of 2.0 degrees per gray shade, add the line:

```
sst_Scaling="-t 1 -s 5 0"
```

in the areas script. This will cover all sea surface temperature values from 0 degrees to 51 degrees. See avhIngest(1) and avhSST(1) for more information.

By default, the following scaling is defined by avhScripts.

channel	type	slope	int	min	max
albedo_ch1	int16	500	0	-65.534	65.534
albedo_ch2	int16	500	0	-65.534	65.534
albedo_ch3	int16	500	0	-65.534	65.534
btemp_ch3	int16	200	200	36.165	363.835
btemp_ch4	int16	200	200	36.165	363.835
btemp_ch5	int16	200	200	36.165	363.835

## Data Base

These variables control information about where the data base of Level-3 and Level-4 will reside and the structure of that data base.

### *Region*

This variable will be used to create the default data base directories. By default it set to *\$MapName*.

### *AvhDataBase*

This variable is used to indicate the location of the image data base for the generated product file. By default, it is set to:

*\$ApsDataBase/\$Level/\$Sensor/\$Version/\$Region/\$Year/\$Month*

where, *\$ApsDataBase* is defined in the *aps.conf* file and represents the top directory of the data base. *\$Level* is set to the string "lv13" by avhInit. *\$Sensor* is set to the string "avhrr" by avhInit. *\$Version* is set to "3.0" by avhInit. *\$Region* is set to "\$MapName" by avhInit. *\$Year* and *\$Month* are set by avhProcess based on the input file.

The user can override *\$AvhDataBase* since it is evaluated by the shell prior to use. For example, if the line:

*AvhDataBase="\\$ApsDataBase/avhrr/\\$Year"*

is set in the areas script and we assume that *\$ApsDataBase* is set to "/data" and that for a particular file *\$Year* has been set to 1999, then the product file will be moved to /data/avhrr/1999. Note that to use the variables, the user must "escape" the '\$' by inserting a '\".

### *AvhDAYDataBase*

This variable is used to indicate the location of the Level-4 daily composites data base for the generated product file. By default, it is set to:

*\$ApsDataBase/\$CompLevel/\$Sensor/\$Version/\$Region/ daily/\$Year/\$Month*

where, *\$ApsDataBase* is defined in the *aps.conf* file and represents the top directory of the data base. *\$CompLevel* is set to the string "lv14" by AvhInit. *\$Sensor* is set to the string "avhrr" by AvhInit. *\$Version* is set to "3.0" by AvhInit. *\$Region* is set to *\$MapName* by avhInit. *\$Year* and *\$Month* are set by *AvhProcess* based on the input file.

The user may override *\$AvhDAYDataBase* since it is evaluated by the shell prior to use.

*AvhNDDataBase*

This variable is used to indicate the location of the Level-4 weekly (8-day) composites data base for the generated product file. By default, it is set to:

*\$ApsDataBase/\$CompLevel/\$Sensor/\$Version/\$Region/weekly/\$Year.*

where, *\$ApsDataBase* is defined in the *aps.conf* file and represents the top directory of the data base. *\$CompLevel* is set to the string "lv14" by AvhInit. *\$Sensor* is set to the string "avhrr" by AvhInit. *\$Version* is set to "3.0" by AvhInit. *\$Region* is set to *\$MapName* by avhInit. *\$Year* is set by *AvhProcess* based on the input file.

The user may override *\$AvhNDDataBase* since it is evaluated by the shell prior to use.

*AvhMODataBase*

This variable is used to indicate the location of the Level-4 monthly composites data base for the generated product file. By default, it is set to:

*\$ApsDataBase/\$CompLevel/\$Sensor/\$Version/\$Region/monthly/\$Year.*

where, *\$ApsDataBase* is defined in the *aps.conf* file and represents the top directory of the data base. *\$CompLevel* is set to the string "lv14" by AvhInit. *\$Sensor* is set to the string "avhrr" by AvhInit. *\$Version* is set to "3.0" by AvhInit. *\$Region* is set to *\$MapName* by avhInit. *\$Year* is set by *AvhProcess* based on the input file.

The user may override *\$AvhMODataBase* since it is evaluated by the shell prior to use.

*AvhYRDataBase*

This variable is used to indicate the location of the Level-4 yearly composites data base for the generated product file. By default, it is set to:

*\$ApsDataBase/\$CompLevel/\$Sensor/\$Version/\$Region/yearly.*

where, *\$ApsDataBase* is defined in the *aps.conf* file and represents the top directory of the data base. *\$CompLevel* is set to the string "lv14" by AvhInit. *\$Sensor* is set to the string "avhrr" by AvhInit. *\$Version* is set to "3.0" by AvhInit. *\$Region* is set to *\$MapName* by avhInit.

The user may override *\$AvhYRDataBase* since it is evaluated by the shell prior to use.

*CmpOpt*

This can be defined by the user to select the type of compression program to call for the output product file before it is moved to *\$AvhDataBase*. This option can be set to: "gzip", "compress", "bzip2" or "none". Only set *CmpOpt* to a compression type that is available on user's machine. *Note:* If the user has defined *XNumChunks* or that variable is defined by avhInit, then the HDF file will automatically be internally compressed and this variable will have no effect.

*XNumChunks, YNumChunks*

These are used to define the number of "chunks" in each direction which will be created by imgReformat program for the product files. By default, a large level-3 file will be rewritten in chunks (or "tiles") with each chunk being compressed. A chunk will be no larger than 640 by 640. So, if the map image is 2430 by 1810, then imgReformat will create a total of 12 chunks (4 across and 3 down). A 1500 by 1500 map image will be divided into 9 chunks (3 across and 3 down). A 600 by 300 map image will NOT be chunked.

*avhVerbose*

If defined this variable will cause the script functions to call 'set -x' within each script function. This will have the effect of printing out each step as it is executed.

*MapFile*

Name of file containing image map file. Defaults to *\$ApsData/maps.hdf*

*MapExt*

This is a string that is appended to the Level-3 file which is written to the database. Usually it is a three character extension all uppercase.

*MinPixels, MinLines*

Used to set the minimum pixels/lines that must be extracted from the AVHRR file by avhExtract to continue processing. These are used to insure that enough of the input file covers the area of interest. By default these are not defined and, therefore, no check is performed.

*AvhAreaOpts*

This allows the use of options to the avhArea(1) program to be added. However, this string should not contain the -p, -l, or -M options.

## Browse Image Variables

*L3BrowseList*

This is a list of whitespace delimited products which are converted to browse images. The products in this list must also be present in the *\$L3ProdList* variable. By default, no browse images are created. That is, BrowseList is not defined.

*L3BrowseDir*

This variable is used to indicate the location of for the browse images. By default, it is set to: *\$ApsImagBase/\$Level/\$Sensor/\$Version/\$Region/\$Year/\$Month* where, *\$ApsImagBase* is set in the *aps.conf* file and represents the top directory of the browse data base. *\$Level* is set to the string "lv13" by avhInit. *\$Sensor* is set to the string "avhrr" by avhInit. *\$Version* is set to "3.0" by avhInit. *\$Region* is set to *\$MapName* by avhInit. *\$Year* and *\$Month* are set by avhProcess based on the input file.

The user can override *\$BrowseDataBase* since it is evaluated by the shell prior to use. For example, if the line:

```
L3BrowseDir="\$ApsImagBase/browse/\$Year"
```

is set in the areas script and we assume that \$ApsImagBase is set to "/data" and that for a particular file \$Year has been set to 1999, then the browse image will be moved to /data/browse/1999.

*\${prod}\_BrowseScaling*

This option is currently only defined for albedo\_ch[1,2], btemp\_ch[3-5], sst, and c\_660 products. It defines the user's desired output scaling for the browse images. For example, to produce an SST browse image that is 300 pixels by 400 lines and has a data range from 0.0 to 30.5 degrees, add the line:

```
sst_BrowseScaling="-r 0.0,30.5 -R 20,199 -s 300,400"
```

in the areas script. See imgBrowse(1) for more information. *Note:* The output range is set to 20 and 199 because the avhMake-Browse file script uses NSIPS files for overlays and colortables by default.

By default, the following scaling is defined by avhScripts.

channel	function	min	max
albedo_ch1	linear	0.0	50.0
albedo_ch2	linear	0.0	50.0
albedo_ch3	linear	0.0	50.0
btemp_ch3	linear	260.0	310.0
btemp_ch4	linear	260.0	310.0
btemp_ch5	linear	260.0	310.0
sst	linear	2.0	35.0
c_660	log10	1.49	50.0

*Note:* For the other products ("K\_PAR", "suspend", "ref\_dif", "Ray\_ch1"), the user must provide the corresponding variables (K\_PAR\_BrowseScaling, suspend\_BrowseScaling, refdif\_BrowseScaling, Ray\_ch1\_BrowseScaling) as there are no defaults.

## Program Variables

These variables define the programs used by avhScripts. The user can override these to test a new version of a program. They are defined by avhInit.

*ApsInfo*

Set to the name of the satellite specific program used to obtain information from input Level-1 file. Defaults to \$ApsBin/avhInfo.

*AvhArea*

Set to the name of the program used to determine if a AVHRR file covers a map. Defaults to \$ApsBin/avhArea.

*AvhInfo*

Set to the name of the program used to obtain information from AVHRR file. Defaults to \$ApsBin/avhInfo.

*AvhClouds*

Set to the name of the program used to produce clouds masks from input calibrated AVHRR images. Defaults to \$ApsBin/avhClouds.

*AvhIngest*

Set to the name of the program used to ingest and calibrate the AVHRR data. Defaults to \$ApsBin/avhIngest.

*AvhSST*

Set to the name of the program used to produce sea surface temperature estimates from input brightness temperature images. Defaults to \$ApsBin/avhSST.

*AvhTurbid*

Set to the name of the program used to produce visible data products from input percent albedo images. Defaults to \$ApsBin/avhTurbid.

## SCRIPT FUNCTIONS

This section describes each script function available.

### ***avhClouds***

This script function run the program avhClouds:

```
$ avhClouds levelsL2.spk levels3.spk levels4.spk levels5.spk levels1.spk levels2.spk
```

If avhClouds returns successfully, the program spkToSDS is called to append the levelsL2.spk file to the Level-2 HDF file. The term 'l2\_flags' is appended to an internal variable used to maintain a list of products which need to be warped.

### ***avhComposites***

#### ***avhDataBase***

This script function checks to see if the variable \$BrowseList is defined. If so, \$BrowseDir is checked for existence and set to a default value if not. For each product in BrowseList, apsAddToWWW is called to move the browse image to the BrowseDir.

Next the script function checks to see if \$L3ProdList is defined. If so, the AvhDataBase variable is checked for existence and set to a default if not. If the user "XNumChunks" was set either by the user or avhInit, then the compression command is set to none since these products are already compressed. Otherwise, the compression is set to \$L3CmpOpt, which may have been previously defined by user using CmpOpt.

### ***avhDefaultBrowseParameters***

This script function sets up the default scaling parameters for the browse images. To begin the function checks to see if \$ImgSamples is defined. If not, then the default image sizes are calculated using imgCalculateSize. For each product in the BrowseList, if the variable \${prod}\_BrowseScaling is not defined it is defined. See BrowseList above for more information and default variables.

### ***avhExtract***

This script function calls `avhArea` to determine if the input AVHRR file covers the map. It also uses the `$MinPixels` and `$MinLines` to determine coverage if they have been set. It returns a 0 if file covers area; a 1 for a processing error; and a 2 if file does not cover area.

## ***avhInit***

This script function defines programs used for AVHRR processing, environmental variables needed by those programs, default browse image parameters, and other miscellaneous variables. It must be called before any other scripts defined here.

## ***avhMakeBrowse***

This script function builds the default browse images. For each product in the `L3BrowseList` variable, the program `imgBrowse` will generate a quick-look image as a JPEG formatted file.

## ***avhMap***

This script function will call the `imgMap` program to warp each image and then call `imgReformat` to "chunkify" it if `XNumChunks` has been defined. Finally, all the attributes from the input HDF Level-2 file are copied over to the HDF Level-3 file.

## ***avhIngest***

This script function will call the `avhIngest` program to produce the required calibrated data that covers the area of interest. For each product in `$L3ProdList`, the data is appended to a HDF Level-2 file.

## ***avhIngestArgs***

This script function builds up the argument string needed by `avhIngest` to output the desired products. It sets up the default scaling.

## ***avhProcess***

This script function is the default processing flow which is described at the top of this man page. It generally is just a series of calls to other script functions defined in this section.

## ***avhProducts***

This script function uses the `$L3ProdList` to determine what channels must be output by `avhIngest`. Some of these outputs may be temporary.

## ***avhProductIngest***

This script function is used by `avhProducts`. It removes duplicates on a list and handles non-calibrated vs calibrated output.

## ***avhSetNames***

This script function defines the standard naming scheme. By default, the HDF Level-2 file will be name: `"$SatCode$Year$DoY$Time.L2_HNAV_$.MapExt"`. *Note:* `MapExt` is defined by the user and the others are defined by the input file and the `avhInfo` program.

## ***avhSST***

This script function calls the program avhSST, if "sst" is defined in the *\$L3ProdList*. If successful, the product is added to the HDF Level-2 file.

### ***avhTurbid***

This script function calls the program avhTurbid, if any of the visible data products ("Ray\_ch1", "ref\_dif", "K\_PAR", "suspend", or "c\_660" is defined in *\$L3ProdList*. If successful, the products are added to the HDF Level-2 file.

### ***avhWarp***

This script function calls the program imgMap to warp the data.

## **CONFORMANCE**

These script function attempt to conform to the IEEE 1003.2 POSIX Shell Standard. They were, however, developed and tested using the Bourne Shell and Korn Shells under IRIX 5.3 and IRIX 6.5 respectively.

---

## Name

avhSwapL0 -- converts a Terascan Level-0 formatted file to NESDIS Level-0 formatted file

avhSwapL0

avhSwapL0 *input.l0 output.l0*

## Description

This program will take an AVHRR "hrpt" Terascan Level-0 formatted file and convert it to a NESDIS Level-0 format required for input to avhL1bgen. The Terascan Level-0 format can be created by the **archive** command of the Terascan system. The input file consists of 44 512-byte sectors in which the 10-bit AVHRR Level-0 stream is merged into 16-bits. This program will swap the bytes so that they resemble the HPRT Minor Frame format described in NOAA Technical Memorandum NESS 107 - Rev 1.

---

# Name

avhTurbid -- produces beam attenuation, diffuse attenuation, suspended solids images from percent albedo images.

avhTurbid

```
avhTurbid [-g angle] [-n angle] input.hdf output.hdf
```

# Description

This program is used to estimate the beam attenuation coefficient (c<sub>660</sub>), the diffuse attenuation coefficient for photosynthetically active radiation for 400-700 nm (KPAR), and total suspended solids (or seston) in coastal waters using the percent albedo for channels 1 and 2 of the AVHRR data. The percent albedo channels are products of the program avhIngest.

The input file must contain the `albedo_ch1` and `albedo_ch2` parameters in the input file from which these products will be generated. images for channels 1 and 2, respectively, before the '.' in the filename. For example, if the first argument is "albedo\_ch.spk", then the file "albedo\_ch1.spk" will hold the percent albedo image for channel 1 and the file "albedo\_ch2.spk" will hold the percent albedo image for channel 2.

Likewise, for the output file name. In this case, the output numbers follow the following table:

Num	Slope	Intercept	Type	Product
1	0.1	0.0	byte	Rayleigh reflectance for channel 1
2	0.05	-0.05	byte	Water reflectance for channel 1
3	1.0	0.0	byte	Total Suspended Solids
4	0.1	0.0	byte	Diffuse Attenuation Coefficient for PAR
5	0.005	0.0	short	Beam Attenuation Coefficient at 660 nm

Therefore, if the second argument is "turbid.spk", then "turbid5.spk" will be an image of the beam attenuation coefficient.

The slope and intercept in the tables will take the integer "gray" value and convert them into the geophysical values. The first four products are stored as 8-bit unsigned integers and the Beam Attenuation product is stored as 16-bit signed integers. Currently, there is no option to change the data conversion.

# Options

-g angle

This sets the glint angle used to determine if an area is susceptible to glint. If an area is found to be glint contaminated, it is masked out. By default, this angle is set to -1.0 to indicate that no glint masking is being performed.

- n angle  
 This is the angle used to determine if for the selected control point, the sun is too low or down. If the solar zenith angle is greater than this angle the control is not processed. By default the angle is set to -1.0.
- help  
 Display program help.
- version  
 Display program name version and time of compilation.

## Files

- nmcoeff.YYYY.d  
 There are a series of files for each year. The files contain the tau for rayleigh and ozone for satellite NOAA-9, NOAA-10, NOAA-11, NOAA-12, NOAA-14, and NOAA-15, NOAA-16.

## Environment Variables

- COEFFDIR  
 This environmental variable points to the location of the coefficients files (nmcoeff.YYYY.d).

## Examples

This example will produce floating point outputs of albedo channels 1 and 2 for input to the turbidity program avhTurbid(1).

### Example 6. Use of avhTurbid

```
$ avhIngest -1 R -2 R NJ1999360192854.L1B_HNAV avhrr.spk
$ avhTurbid avhrr.spk turbid.spk
Turbid 0
$ ls -l turbid*.spk
-rw-r----- 1 aps  aps  4168192 Jan  6 12:37 turbid1.spk
-rw-r----- 1 aps  aps  4168192 Jan  6 12:37 turbid2.spk
-rw-r----- 1 aps  aps  4168192 Jan  6 12:37 turbid3.spk
-rw-r----- 1 aps  aps  4168192 Jan  6 12:37 turbid4.spk
-rw-r----- 1 aps  aps  8335872 Jan  6 12:37 turbid5.spk
```

Notice that by default, turbid5.spk, is written out as signed 16-bit integers. This product is the beam attenuation at 660 nm (c\_660).

## References

---

Stumpf, R. P. (1992) *Remote Sensing of Water Clarity and Suspended Sediments in Coastal Waters*. In Proceedings of the First Thematic Conference on Remote Sensing for Marine and Coastal Environments. SPIE 1930, 293-305. Environmental Research Institute of Michigan, Ann Arbor, MI.

Gould, R. W. and Arnone, R. A. (1997) *Estimating the beam attenuation coefficient in coastal waters from AVHRR imagery*. Continental Shelf Research, Vol 17, No 11, p 1375-1387.